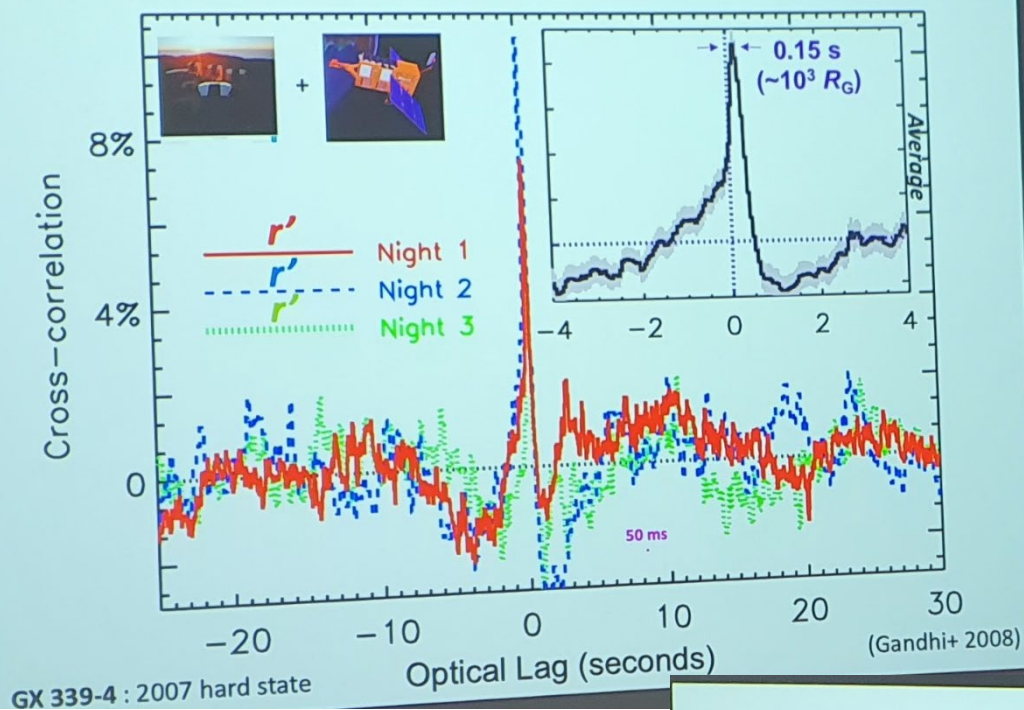


# 30<sup>th</sup> Texas Symposium

# Sub-second X-O Cross Correlation Function (CCF)



# Origin of sub-second delays?

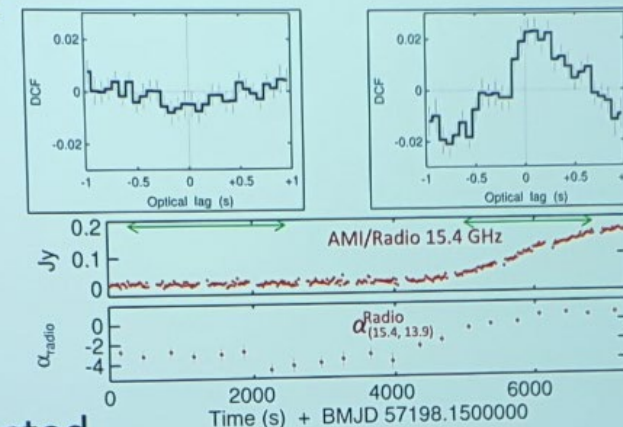
~~Reprocessing?~~  
No. Too short.  
Too red.

• Jet?

~~Hot flow?~~

No radio expected.

No opt./X anti-correlation. No QPOs.



# Answering the Open Questions

1) What are the physical conditions associated with jet formation and launching in XRBs?

⇒ *B field ~ 10<sup>4</sup> G. Inner jet zone size ~0.1 light-second, but variable (what does it depend on?)*

2) Can we see the accretion-ejection connection in *real* time?

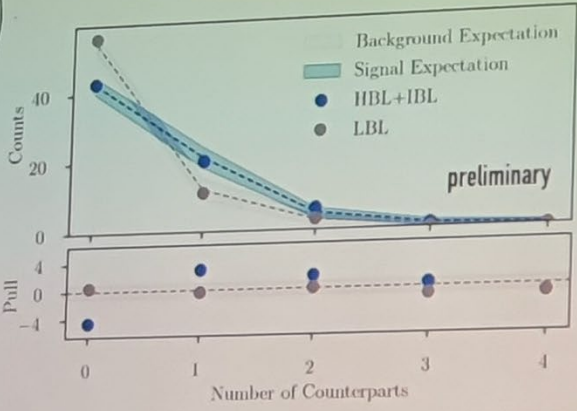
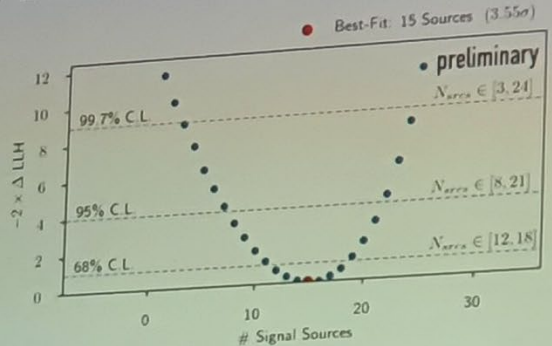
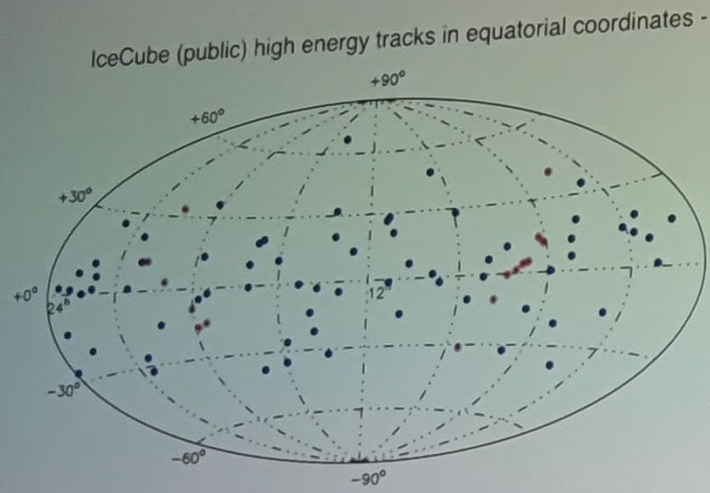
⇒ *Yes. Sub-second multiwavelength lags explained by internal shocks.*

3) Can we tie models (MHD) with observational constraints (radiation)?

⇒ *Evidence for a universal standing shock at ~10<sup>3</sup> R<sub>G</sub>. Does this apply to AGN?*

## 6) Evidence of neutrinos from blazars ... growing

P. Glommi, T. Glauch, P. Padovani, E. R., A. Turcati, Y.L. Chang Dissecting the regions around IceCube high-energy neutrinos: growing evidence for the blazar connection, submitted to MNRAS

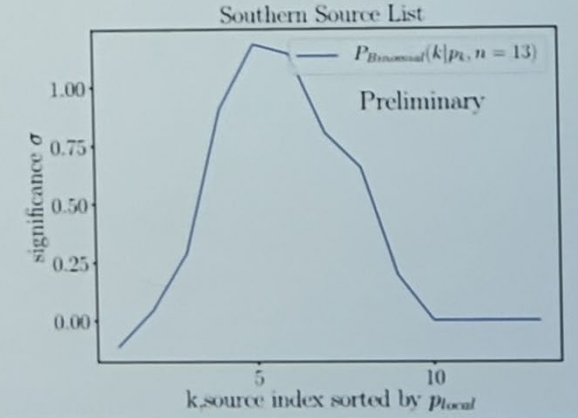
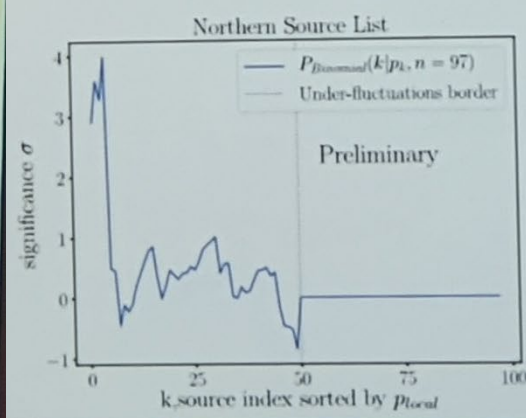


More than 70 gamma-ray emitting blazars found in spatial coincidence

# Source Population Results

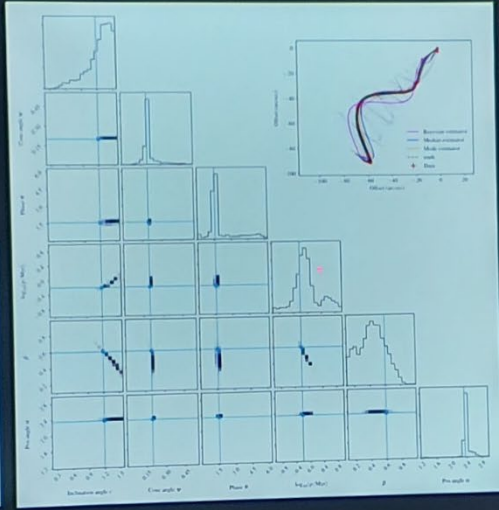
Search for excess of hotspots → A significant p-value demonstrates inconsistency with background-only for entire catalog.

- Probability of k or more sources passing a threshold out of a catalog of N.
- 4σ pre-trial where k=4 in Northern Catalog.  
→ 3.3σ post-trial. (2.25σ w/o TXS 0506+056) to account for N other possible excesses
- Includes NGC 1068, TXS 0506+056, PKS 1424+240, GB6 J1542+6129



# MCMC SIMULATIONS

- \* Parameter searches<sup>1</sup> on single and multiple jets
- \* Good constraints on precession period, phase, cone opening and position angle



## A Markov Chain Monte Carlo approach for measurement of jet precession in radio-loud active galactic nuclei

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Accepted XXX, Received YYY, in original form ZZZ

### ABSTRACT

Jet precession can reveal the presence of binary systems of supermassive black holes. The jet precession can be measured by monitoring the apparent position of the jet on the sky. This is an important challenge in the context of constraining the binary supermassive black hole formation rate, which are expected as a result of hierarchical galaxy evolution. The jet's size, position, and orientation along the line of sight of a given source often result in uncertainties regarding jet path. This paper presents a new approach for the determination of precession parameters using a 2D MCMC curve-fitting algorithm. Applying the method to Cygnus A, we find evidence for precession, suggesting that the source is precessing. Inferred in the context of binary black holes leads to a constraint of period, size and likely sub-parsec orbital separation for the putative supermassive binary.

**Key words:** galaxies: active – galaxies: jets – radio continuum: galaxies – methods: data analysis – methods: statistical

# CYGNUS A

Real universe example

Cygnus A: large well-studied radio source showing precession characteristics and multiple visible knots in jet and counterjet



